# Exponential integrators for the Schrödinger equation with time-dependent Hamiltonian

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- 2 H(t) is a real symmetric matrix

## The Goal

 The numerical integration of the time-dependent Schrödinger Equationt (ħ = 1)

$$i\frac{\partial}{\partial t}\psi(x,t) = -\frac{1}{2\mu}\nabla^2\psi(x,t) + V(x,t)\psi(x,t)$$

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 A quantum two-level system can be written down in the form

$$H(t) = \left( egin{array}{cc} \omega(t) & \mathcal{C}(t) \ \mathcal{C}^*(t) & -\omega(t) \end{array} 
ight)$$

where  $\omega(t)$  is a real function and C(t) is, in general, a complex function of *t*.

## Frequently used method

The exponential midpoint method to advance from  $t_n$  to  $t_{n+1} = t_n + h$ 

$$u_{n+1} = e^{-ihH(t_n+\frac{h}{2})}u_n$$

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More efficient methods to reach high accuracy?

## Methods for the non-autonomous case

## I) H(t) is a complex Hermitian matrix



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Peskin, Kosloff and Moiseyev (93-94)

$$i\frac{\partial}{\partial t}\psi(\mathbf{x},t',t)=\mathcal{H}(\mathbf{x},t')\psi(\mathbf{x},t',t),$$

where

$$\mathcal{H}(\mathbf{x},t') = \hat{H}(\mathbf{x},t') - i\frac{\partial}{\partial t'}$$

and then standard methods for the autonomous case can be used with very large time steps.

Main trouble: the linear system to be solved is of a higher dimension, making the algorithms computationally very costly.

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$$i\frac{d}{dt}u = H(t)u$$

is equivalent to  $(U = (u, u_t) \in \mathbb{C}^{d+1})$ 

$$U' = \frac{d}{dt} \left\{ \begin{array}{c} u \\ u_t \end{array} \right\} = \left\{ \begin{array}{c} -iH(u_t) u \\ 0 \end{array} \right\} + \left\{ \begin{array}{c} 0 \\ 1 \end{array} \right\} = A(U) + B(U)$$

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$$u_{n+1} = \prod_{j=1}^{m} e^{-i a_j h H(t_n + c_j h)} u_n,$$

 $\{a_j, b_j\}_{j=1}^m$  splitting method  $(c_j = \sum_{k=1}^j b_k)$ . Efficient methods: 4th-order (m = 6); 6th-order (m = 10).

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$$\Omega^{[4]} = B_1 + [B_2, B_3],$$
 with  $B_i = h \sum_{j=1}^K a_{i,j} A(t_n + c_j \tau)$ 

 $(\Omega^{[r]} = \Omega + \mathcal{O}(h^{r+1}))$  so that  $\Omega^{[4]}u = v_1 + (v_4 - v_5)$  with

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 $\Omega^{[6]}u$  involves at least 13 products.

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- m = 3: order 5 (with complex coefficients  $a_{ik}$ ).

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m = 5: order 6 with one commutator ( $[B_{3,1}, B_{3,2}] = \mathcal{O}(h^3)$ )

$$u_{n+1} = e^{B_5} e^{B_4} e^{[B_{3,1},B_{3,2}]} e^{B_2} e^{B_1} u_n$$

(real **POSTIVE** coefficients  $a_{ik}$ ).

#### The Rosen–Zener model

The Hamiltonian in terms of Pauli matrices is given by

$$H(t) = rac{1}{2}\omega\sigma_3 + V(t)\sigma_1$$

and in the interaction picture

$$H_{l}(t) = \cos(\omega t) V(t) \sigma_{1} - \sin(\omega t) V(t) \sigma_{2} \in \mathbb{C}^{2 \times 2}$$

A Rosen–Zener model with dissipation of dimension n = 2k:

 $H_{l}(t) = f_{1}(t) \sigma_{1} \otimes I_{k} + f_{2}(t) \sigma_{2} \otimes R_{k} + \delta D_{n}.$ 

with 
$$D_n = i \times \text{diag}\{-1, -4, \dots, -n^2\}, R_n = \text{tridiag}\{1, 0, 1\},$$
  
 $f_1(t) = \frac{V_0 \cos(\omega t)}{\cosh(t/T)}, \quad f_2(t) = -\frac{V_0 \sin(\omega t)}{\cosh(t/T)}, \quad \delta > 0,$   
 $T = 1, t \in [-5T, 5T], n = 10.$ 

Example 1:



solid lines: 3-exp 5th-order with complex coefs. dashed lines: 6-exp 6th-order with real coefs. Taylor of order 4 (circles), 6 (squares) and 8 (stars)

## Example 1:



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## Example 1: with dissipation



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## II:*H*(*t*) is a real symmetric matrix



## The autonomous case

$$i \frac{d}{dt} u = H u \quad \Rightarrow \quad u(T) = e^{-iTH}u(0)$$

where  $u \in \mathbb{C}^d$  and  $H \in \mathbb{R}^{d \times d}$  is a real and symmetric matrix.

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where  $u \in \mathbb{C}^d$  and  $H \in \mathbb{R}^{d \times d}$  is a real and symmetric matrix. Formally, the problem to solve is

$$i\frac{du}{dt} = P^{-1} \begin{pmatrix} E_0 & & \\ & E_1 & \\ & & \ddots & \\ & & & E_{d-1} \end{pmatrix} Pu = Hu$$

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which is just a set of *d* harmonic oscillators

$$\left\{\begin{array}{c} q_T\\ p_T\end{array}\right\} = \left(\begin{array}{cc} T_1^m & T_2^m\\ -T_2^m & T_1^m\end{array}\right) \left\{\begin{array}{c} q_0\\ p_0\end{array}\right\}$$

$$\frac{T\beta}{m} < 0.37$$

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Chebyshev method

$$\left\{ \begin{array}{c} q_C \\ p_C \end{array} \right\} = \left( \begin{array}{cc} C_1^m & C_2^m \\ -C_2^m & C_1^m \end{array} \right) \left\{ \begin{array}{c} q_0 \\ p_0 \end{array} \right\}$$

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$$\frac{T\beta}{m} < 0.90$$

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Symplectic methods

$$\left(\begin{array}{cc}1&0\\-b_ky&1\end{array}\right)\left(\begin{array}{cc}1&a_ky\\0&1\end{array}\right)=\left(\begin{array}{cc}1&a_ky\\-b_ky&1-a_kb_ky^2\end{array}\right)$$

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$$K(y) \equiv \prod_{k=1}^{m} \begin{pmatrix} 1 & a_k y \\ -b_k y & 1 - a_k b_k y^2 \end{pmatrix} = \begin{pmatrix} K_1^{2m-2} & K_2^{2m-1} \\ K_3^{2m-1} & K_4^{2m} \end{pmatrix}$$

$$\left\{ \begin{array}{c} q_{S} \\ p_{S} \end{array} \right\} = \left( \begin{array}{cc} K_{1}^{2m-2} & K_{2}^{2m-1} \\ K_{3}^{2m-1} & K_{4}^{2m} \end{array} \right) \left\{ \begin{array}{c} q_{0} \\ p_{0} \end{array} \right\} \qquad \frac{T\beta}{m} < 2$$

#### Horner's algorithm for the Taylor method:

$$y_0 = u_0$$
  
**do**  $k = 1, m$   

$$y_k = u_0 - i \frac{T\beta}{m+1-k} \tilde{H} y_{k-1}$$
  
**enddo**  

$$w_T = y_m$$

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 $w_T = y_m$ 

Clenshaw algorithm for Chebyshev ( $c_k = (-i)^k J_k(T \beta)$ ):

$$d_{m+2} = 0, \quad d_{m+1} = 0$$
  
do  $k = m, 0$   
 $d_k = c_k u_0 + 2\tilde{H}d_{k+1} - d_{k+2}$   
enddo  
 $w_C \equiv P_{m-1}^C(T\beta\tilde{H}) u_0 = d_0 - d_2$ 

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Splitting Symplectic methods:

do 
$$k = 1, m$$
  
 $q := q + a_k T \beta \tilde{H} p$   
 $p := p - b_k T \beta \tilde{H} q$   
enddo

$M_m^{(\theta/m)}$	m	$\beta \tau_{\rm max}$	y <sub>*</sub> / m	$\epsilon( heta)$	$\mu( heta)$	$\nu(\theta)$
$M_{10}^{(0.5)}$	10	5	0.63	$3.6 imes10^{-8}$	$8.7 \times 10^{-11}$	$9.8 imes10^{-8}$
M <sub>10</sub> <sup>(0.9)</sup>	10	9	0.94	$3.4 imes10^{-5}$	$2.9 imes10^{-5}$	$1.1 \times 10^{-5}$
$M_{20}^{(0.6)}$	20	12	0.79	$1.6  imes 10^{-13}$	$1.4  imes 10^{-13}$	$5.8  imes 10^{-14}$
$M_{20}^{(1)}$	20	20	1.1	$4.1 \times 10^{-7}$	$1.8  imes 10^{-8}$	$4.8  imes 10^{-7}$
$M_{30}^{(0.75)}$	30	22.5	0.84	$8.1 \times 10^{-15}$	$3.3  imes 10^{-16}$	$1.5  imes 10^{-14}$
$M_{30}^{(1)}$	30	30	1.0	$4.1  imes 10^{-10}$	$1.9 \times 10^{-10}$	$3.1 \times 10^{-10}$
$M_{30}^{(1.3)}$	30	39	1.36	$2.3\times10^{-5}$	$5.2 imes10^{-6}$	$2.2  imes 10^{-5}$
$M_{40}^{(1)}$	40	40	1.1	$1.8  imes 10^{-12}$	$4.9 \times 10^{-14}$	$2.4  imes 10^{-12}$
M <sub>40</sub> <sup>(1.2)</sup>	40	48	1.26	$2.1  imes 10^{-8}$	$2.1  imes 10^{-8}$	$5.3  imes 10^{-10}$
$M_{40}^{(1.4)}$	40	56	1.48	$1.48  imes 10^{-5}$	$4.0  imes 10^{-6}$	$1.7  imes 10^{-5}$
$M_{50}^{(1)}$	50	50	1.07	$4.5\times10^{-15}$	$4.5  imes 10^{-15}$	$2.0 \times 10^{-17}$
$M_{50}^{(1.1)}$	50	55	1.13	$4.5 imes10^{-13}$	$4.2  imes 10^{-13}$	$4.1 \times 10^{-14}$
$M_{50}^{(1.2)}$	50	60	1.26	$5.4  imes 10^{-11}$	$2.7 \times 10^{-11}$	$3.8 \times 10^{-11}$
$M_{50}^{(1.3)a}$	50	65	1.32	$1.2  imes 10^{-8}$	$1.2  imes 10^{-8}$	$8.3  imes 10^{-10}$
$M_{50}^{(1.3)b}$	50	65	1.32	$5.9  imes 10^{-7}$	$9.5  imes 10^{-11}$	$6.1 \times 10^{-7}$
$M_{60}^{(1.1)}$	60	66	1.15	$7.2  imes 10^{-15}$	$7.2  imes 10^{-15}$	$2.6  imes 10^{-17}$
$M_{60}^{(1.2)a}$	60	72	1.3	$1.5  imes 10^{-12}$	$1.1 \times 10^{-12}$	$8.3  imes 10^{-13}$
$M_{60}^{(1.2)b}$	60	72	1.26	$4.2 \times 10^{-11}$	$6.5 \times 10^{-14}$	$4.6 \times 10^{-11}$
$M_{60}^{(1.3)}$	60	78	1.36	$1.2  imes 10^{-9}$	$7.8 \times 10^{-11}$	$1.2 \times 10^{-9}$
$M_{60}^{(1.4)a}$	60	84	1.41	$8.4 \times 10^{-8}$	$2.4  imes 10^{-8}$	$7.4 \times 10^{-8}$
$M_{60}^{(1.4)b}$	60	84	1.46	$2.9 imes10^{-6}$	$3.7  imes 10^{-9}$	$2.9  imes 10^{-6}$

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## Example 1:



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## Numerical example 1

(Lubich, Blue book, 2008) To approximate

$$e^{-iH}u_0$$

with  $u_0$  a unitary random vector and

$$H = \frac{\lambda}{2} \begin{pmatrix} 2 & -1 & & \\ -1 & 2 & -1 & & \\ & \ddots & & \\ & & -1 & 2 & -1 \\ & & & -1 & 2 \end{pmatrix} \in \mathbb{R}^{N \times N}, \quad N = 10000$$

 $0 \le E_k \le 2\lambda$ , k = 1, 2, ..., 10000After a shift,  $H - \lambda I$ , we can take:  $T\beta = \lambda = t$ We approximate:

$$e^{-it}e^{-it\hat{H}}u_0, \qquad \hat{H}=(H-tI)/t$$

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## Example:



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## The Mathematical Problem

$$i \frac{d}{dt} u(t) = H(t) u(t), \qquad u(0) = u_0 \in \mathbb{C}^N,$$

It can be written as the 2N-dimensional real system

$$q' = H(t)p, \qquad p' = -H(t)q.$$

Classical Hamiltonian equations associated to the classical Hamiltonian

$$\mathcal{H}(q,p,t) = \frac{1}{2}p^{T}H(t)p + \frac{1}{2}q^{T}H(t)q.$$

Sanz-Serna&Portillo (1996): Time as two dependent variables:

$$\bar{H} = \left(\frac{1}{2}p^{T}H(p_1)p + p_2\right) + \left(\frac{1}{2}q^{T}H(q_2)q - q_1\right) = A(P) + B(Q)$$

with  $q_1, q_2, p_1, p_2 \in \mathbb{R}$ 

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## The Proposed Algorithm

$$q_0 = Re(u_n), \quad p_0 = Im(u_n)$$
  
 $H_1 = H(t_n + c_1 h), \quad H_2 = H(t_n + c_2 h), \quad H_3 = H(t_n + c_3 h)$ 

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do 
$$i = 1, m$$
  
 $v = (a_{i,1}H_1 + a_{i,2}H_2 + a_{i,3}H_3) p_{i-1}$   
 $q_i = q_{i-1} + hv$   
 $v = (b_{i,1}H_1 + b_{i,2}H_2 + b_{i,3}H_3) q_i$   
 $p_i = p_{i-1} - hv$   
enddo

 $u_{n+1} = q_m + ip_m$ 

The system can be written as

$$z' = \begin{pmatrix} 0 & H(t) \\ -H(t) & 0 \end{pmatrix} z = (A(t) + B(t))z$$

where  $z = (q, p)^T$  and

$$A(t) = \left( egin{array}{cc} 0 & H(t) \ 0 & 0 \end{array} 
ight), \qquad B(t) = \left( egin{array}{cc} 0 & 0 \ -H(t) & 0 \end{array} 
ight).$$

The Magnus series expansion allows to write the formal solution in exponential form

$$z(t+h) = e^{\Omega(t,h)}z(t), \qquad \Omega(t,h) = \sum_{k=1}^{\infty} \Omega_k(t,h)$$

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Proposed methods

$$z(t+h) \approx e^{\tilde{A}_{m+1}} e^{\tilde{B}_m} e^{\tilde{A}_m} \cdots e^{\tilde{B}_1} e^{\tilde{A}_1} z(t)$$
  
$$\approx \begin{pmatrix} I & \tilde{H}_{m+1}^A \\ 0 & I \end{pmatrix} \begin{pmatrix} I & 0 \\ -\tilde{H}_m^B & I \end{pmatrix} \cdots \begin{pmatrix} I & 0 \\ -\tilde{H}_1^B & I \end{pmatrix} \begin{pmatrix} I & \tilde{H}_1^A \\ 0 & I \end{pmatrix} z(t)$$

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where

$$\tilde{H}_i^A = h \sum_{j=1}^k a_{i,j} H(t+c_j h), \qquad \tilde{H}_i^B = h \sum_{j=1}^k b_{i,j} H(t+c_j h),$$

for appropriate coefficients  $c_i, a_{i,j}, b_{i,j}$ .

#### SE with Morse Potential and external laser interaction

$$i\frac{\partial}{\partial t}\psi(x,t) = \left(-\frac{1}{2\mu}\frac{\partial^2}{\partial x^2} + V(x) + f(t)x\right)\psi(x,t)$$

with

$$V(x) = D(1 - e^{-\alpha x})^2, \qquad f(t)x = A\cos(\omega t)x$$

 $\mu = 1745$ , D = 0.2251,  $\alpha = 1.1741$  *a.u.* (HF molecule in a.u.), A = 0.011025 and laser frequency  $\omega = 0.01787$ .  $x \in [-0.8, 4.32]$ , split into N = 64 parts and periodic bc. Initial conditions

$$\phi(\mathbf{x}) = \sigma \exp\left(-(\gamma - 1/2)\alpha \mathbf{x}\right) \exp\left(-\gamma e^{-\alpha \mathbf{x}}\right),$$

 $\gamma = 2D/w_0$ ,  $w_0 = \alpha \sqrt{2D/\mu}$  ( $\sigma$  is a normalizing constant).

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Kormann, Holmgren, and Karlsson, J. Chem. Phys. (2008).

## Example 1:



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#### The Rosen–Zener model

 $H(t) = \omega \ \sigma_3 \otimes I_k + f(t) \ \sigma_2 \otimes R_k$ 

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$$R_n = \text{tridiag}\{1, 0, 1\},$$
  
$$f(t) = \frac{V_0}{\cosh(t/T)}, \quad V_0 = 10, \quad \omega = 10$$
  
$$T = 1, \ t \in [-5T, 5T], \ n = 10.$$



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#### **Diffusion-advection-reaction equation**

We consider the equation

 $\partial_t u(x,t) = A(x,t) u(x,t) = (\alpha(x,t) \partial_{xx} + \beta(x,t) \partial_x + \gamma(x,t)) u(x,t)$ 

$$\alpha(\mathbf{x},t) = \mathbf{e}^{-\cos x} (\sin t)^2, \qquad \gamma(\mathbf{x},t) = \mathbf{e}^{\sin x} (\mathbf{1} + \mathbf{e}^{-t}),$$

subject to the initial condition  $u(x,0) = \sin(2x)$  and periodic boundary conditions on the spatial interval  $\Omega = [0, 2\pi]$ . Number of grid points M = 100



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#### For solving the nonlinear perturbed problem

$$\frac{du}{dt} = A(t) u + \varepsilon g(t, u),$$

with  $|\varepsilon| \ll 1$  it can be very useful to have a good integrator for the linear part.

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#### H real and constant

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#### H real and t-dependent

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#### H: t-dependent, complex and/or dissipative

SB, F. Casas, and M. Thalhammer, Work in progress.

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Group webpage: http://www.gicas.uji.es

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## Thank You ( D) ( B) ( E) ( E) E ) ( C)